

# Kepler's Laws

1<sup>st</sup> Law: The planets move about the sun in ELLIPTICAL orbits, with the sun at one focus of the ellipse.

2<sup>nd</sup> Law: The straight line joining the sun and a given planet sweeps \_\_\_\_\_  
\_\_\_\_\_.

→ Can be remembered as \_\_\_\_\_

3rd Law: The square of the period of revolution of a planet ABOUT THE SUN is proportional to the cube of its mean distance from the sun.

→ Stated in equation form as \_\_\_\_\_

Defining the variables:

T = \_\_\_\_\_

R = \_\_\_\_\_

r = \_\_\_\_\_

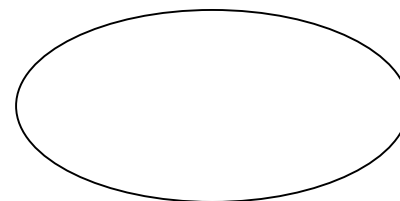
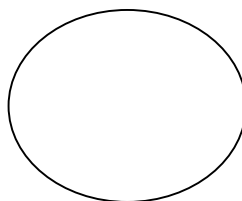
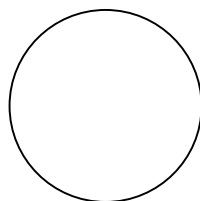
## Some key things to remember/know about Kepler's Laws

### 1<sup>st</sup> Law:

- ✓ Circles have centers. Ellipses are like flattened circles, that don't have a center, but rather have two \_\_\_\_\_.
- ✓ **Eccentricity** may be interpreted as a measure of how much an orbit's shape deviates from a circle.

For a circle,  $e = 0$

For an ellipse,  $0 < e < 1$  (the lower the e value, the more circular the orbit)



- ✓  $e = 0.017$  for Earth's orbit,  $e = 0.093$  for Mars' orbit,  $e = .252$  for Pluto's orbit

## 2<sup>nd</sup> Law:

- ✓ Planets move \_\_\_\_\_ when they are on the side of their elliptical orbit that is closest to the sun.
- ✓ Between March 21 and September 21, there are three days more than between September 21 and March 21. These two dates are the spring and fall equinoxes, when the days and nights are of equal length. Between the equinoxes, the Earth moves  $180^\circ$  around its orbit with respect to the sun. Using Kepler's 2<sup>nd</sup> Law, explain clearly how you can determine the part of the year during which the Earth is closer to the sun.

## 3<sup>rd</sup> Law:

- ✓ All planets that orbit the sun have the same Kepler Constant (which equals \_\_\_\_\_ )
- ✓ All "things" (little) that orbit the same "THING" (BIG) have the same Kepler constant.
- ✓ The orbital period of the earth about the sun is approximately \_\_\_\_\_.
- ✓ The orbital period of the moon about the earth is approximately \_\_\_\_\_.
- ✓ When using Kepler's 3<sup>rd</sup> Law, make sure to use units of METERS and SECONDS.
  
- ✓ The units of K are \_\_\_\_\_.
  
- ✓ When using your calculator with BIG numbers that involve exponents, make sure to utilize parenthesis properly, making sure to pay attention to the ORDER OF OPERATIONS (remember PEMDAS)
  
- ✓ If you solve Kepler's 3<sup>rd</sup> Law for R, it will involve a **cube root**. There are two ways to do this on your calculator.

**Accelerated Physics**  
**“Kepler’s Laws” Worksheet**



1. Using the table below, find the Kepler Constant for each of the objects below (including the moon, but excluding the sun). Explain why the answers make sense.

Object	Mass (kg)	Radius of object (m)	Period of rotation on axis (s)	Mean radius of orbit (m)	Period of revolution of orbit (s)	Kepler constant $R^3/T^2$ ( $m^3/s^2$ )
Sun	$1.98 \times 10^{30}$	$6.95 \times 10^8$	$2.14 \times 10^6$	---	---	---
Mercury	$3.28 \times 10^{23}$	$2.57 \times 10^6$	$5.05 \times 10^6$	$5.79 \times 10^{10}$	$7.60 \times 10^6$	
Venus	$4.83 \times 10^{24}$	$6.31 \times 10^6$	$2.1 \times 10^7$	$1.08 \times 10^{11}$	$1.94 \times 10^7$	
Earth	$5.98 \times 10^{24}$	$6.38 \times 10^6$	$8.61 \times 10^4$	$1.49 \times 10^{11}$	$3.16 \times 10^7$	
Mars	$6.37 \times 10^{23}$	$3.43 \times 10^6$	$8.85 \times 10^4$	$2.28 \times 10^{11}$	$5.94 \times 10^7$	
Jupiter	$1.90 \times 10^{27}$	$7.18 \times 10^7$	$3.54 \times 10^4$	$7.78 \times 10^{11}$	$3.74 \times 10^8$	
Saturn	$5.67 \times 10^{26}$	$6.03 \times 10^7$	$3.60 \times 10^4$	$1.43 \times 10^{12}$	$9.30 \times 10^8$	
Uranus	$8.80 \times 10^{25}$	$2.67 \times 10^7$	$3.88 \times 10^4$	$2.87 \times 10^{12}$	$2.66 \times 10^9$	
Neptune	$1.03 \times 10^{26}$	$2.48 \times 10^7$	$5.69 \times 10^6$	$4.50 \times 10^{12}$	$5.20 \times 10^9$	
Pluto	$6 \times 10^{23}$	$3 \times 10^6$	$5.51 \times 10^5$	$5.9 \times 10^{12}$	$7.82 \times 10^9$	
moon	$7.34 \times 10^{22}$	$1.74 \times 10^6$	$2.36 \times 10^6$	$3.8 \times 10^8$	$2.36 \times 10^6$	

- A planet’s mean distance from the sun is  $2.0 \times 10^{11}$  m. What is its orbital period?
- If a small planet were discovered whose orbital period was twice that of the Earth, how many times farther from the sun would this planet be?
- Using the data from the table below, determine the Kepler constant for any satellite of the Earth. (Note: The moon is a satellite of the Earth.)

**Natural Satellites in the Solar System\***

Planet	Mass ( $m_E=1$ )	Satellite	Orbital Radius ( $\times 10^6$ m)	Period (d)
Earth	1.00	Moon	384.4	27.322
Mars	0.107	Phobos	9.38	0.319
		Deimos	23.46	1.262
Jupiter	318	Thebe	221.9	0.675
		Io	421.6	1.769
		Europa	670.9	3.551
		Elara	11,737	259.7
Saturn	95.2	Janis	151.47	0.695
		Mimas	185.54	0.942
		Calypso	294.67	1.888
Uranus	14.6	Miranda	129.4	1.414
		Ariel	191.0	2.520
		Oberon	583.5	13.463
Neptune	16.72	Triton	355.3	5.877
		Nereid	5,510	360.21
Pluto	0.002	Charon	19.7	6.387

\*"Planetary Satellites: An Update", *Sky and Telescope*, November 1983.

5. List the planets (that orbit the sun) in order, starting with the one closest to the sun and leaving Pluto off, since it's not "really" a planet anymore. You might need to use the internet as a resource.
6. According to Kepler's Laws, which planet takes longer to orbit the sun, Saturn or Neptune? Explain.
7. Calculate the Kepler constants by determining the  $R^3/T^2$  values for each of the planet systems listed in the chart at the bottom of the previous page.
8. An asteroid of diameter 100 km has a mean radius of orbit of  $4.8 \times 10^{11}$  m. What will be its orbital period around the sun?
9. A spy satellite is located one Earth radius above the surface of the Earth. What is its period of revolution?
10. Mars has two moons, Phobos and Deimos (Fear and Panic, the companions of Mars, the god of war). Deimos has a period of 30 h 18 min and a mean distance from the centre of Mars of  $2.3 \times 10^4$  km. If the period of Phobos is 7 h 39 min, what mean distance is it from the centre of Mars?
11. A Martian lander is to be placed in orbit around Mars at a mean altitude of 100 km. What will be the period of the Martian lander?
12. Communications satellites are placed in orbit so that they remain stationary relative to a specific area on the Earth's surface. They are given the name **synchronous satellites** because, to maintain such a position, their period as they orbit must be the same as the Earth's. What is the height of such a satellite measured from (a) the centre of the Earth, and (b) the surface of the Earth?

### ANSWERS

1) $\sim 3.35 \times 10^{18} \text{ m}^3/\text{s}^2$ for all except for the moon.	5) in class discussion	10) $9.2 \times 10^6 \text{ m}$
2) $4.9 \times 10^7 \text{ sec}$	6) in class discussion	
	7)	
	Earth: $1.01 \times 10^{13} \text{ m}^3/\text{s}^2$	
	Mars: $1.09 \times 10^{12} \text{ m}^3/\text{s}^2$	
	Jupiter: $3.21 \times 10^{15} \text{ m}^3/\text{s}^2$	
	Saturn: $9.64 \times 10^{14} \text{ m}^3/\text{s}^2$	
	Uranus: $1.45 \times 10^{14} \text{ m}^3/\text{s}^2$	
	Neptune: $1.74 \times 10^{14} \text{ m}^3/\text{s}^2$	
	Pluto: $2.51 \times 10^{10} \text{ m}^3/\text{s}^2$	
3) 1.6 earth radii	8) $1.82 \times 10^8 \text{ sec}$	11) 6,352.6 sec
4) $1.01 \times 10^{13} \text{ m}^3/\text{s}^2$	9) $1.43 \times 10^4 \text{ sec}$	12) $4.2 \times 10^7 \text{ m}$ ; $3.6 \times 10^7 \text{ m}$